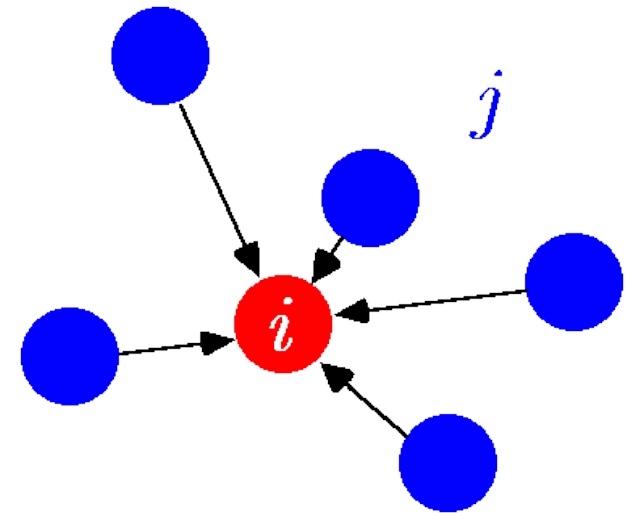


- hard spheres etc. – collisions
- “classical” MD – integration of the equations of motion
- Brownian (stochastic) dynamics, dissipative particle dynamics = MD + random forces

Forces are needed:

$$\vec{f}_i = -\frac{\partial U(\vec{r}^N)}{\partial \vec{r}_i} \quad i = 1, \dots, N$$



Example – pair forces:

$$U = \sum_{i < j} u(r_{ij}) \quad \Rightarrow \quad \vec{f}_i = \sum_{\substack{j=1 \\ j \neq i}}^N \vec{f}_{ji} \equiv - \sum_{\substack{j=1 \\ j \neq i}}^N \frac{du(r_{ji})}{dr_{ji}} \frac{\partial r_{ji}}{\partial \vec{r}_i} = - \sum_{\substack{j=1 \\ j \neq i}}^N \frac{du(r_{ji})}{dr_{ji}} \frac{\vec{r}_{ji}}{r_{ji}}$$

Notation: $\vec{r}_{ij} = \vec{r}_j - \vec{r}_i$, $r_{ij} = |\vec{r}_{ij}|$

$$\frac{d^2 \vec{r}_i}{dt^2} = \ddot{\vec{r}}_i = \frac{\vec{f}_i}{m_i}, \quad i = 1, \dots, N$$

Method of finite differences, timestep h

Initial value problem: \vec{r} and $\dot{\vec{r}}$ at time t_0 are known

Methods:

- Runge–Kutta: many evaluations of the right-hand side/step (costly!)
- Predictor–corrector: a bit better, rarely used
- Verlet and clones (symplectic = good energy conservation)
- Multiple timestep methods: more timescales (usually symplectic)
- Geometric integrators (symplectic)

Taylor expansion:

$$\begin{aligned}\vec{r}_i(t-h) &= \vec{r}_i(t) - h\dot{\vec{r}}_i(t) + \frac{h^2}{2}\ddot{\vec{r}}_i(t) - \dots && +1\times \\ \vec{r}_i(t) &= \vec{r}_i(t) && -2\times \\ \vec{r}_i(t+h) &= \vec{r}_i(t) + h\dot{\vec{r}}_i(t) + \frac{h^2}{2}\ddot{\vec{r}}_i(t) + \dots && +1\times\end{aligned}$$

\Rightarrow numeric 2nd derivative: $\ddot{\vec{r}}_i(t) = \frac{\vec{f}_i(t)}{m_i} = \frac{\vec{r}_i(t-h) - 2\vec{r}_i(t) + \vec{r}_i(t+h)}{h^2} + \mathcal{O}(h^2)$

Verlet method: $\vec{r}_i(t+h) = 2\vec{r}_i(t) - \vec{r}_i(t-h) + h^2 \frac{\vec{f}_i(t)}{m_i}$

Initial values: $\vec{r}_i(t_0-h) = \vec{r}_i(t_0) - h\dot{\vec{r}}_i(t_0) + \frac{h^2}{2} \frac{\vec{f}_i(t_0)}{m_i} + \mathcal{O}(h^3)$

⊕ time-reversible (\Rightarrow no energy drift); even symplectic

⊖ cannot use for $\ddot{r} = f(r, \dot{r})$ because $\dot{r}(t)$ is not known at time t

Identical trajectories: leap-frog, velocity Verlet, Gear ($m = 3$), Beeman

Leap-frog

velocity = displacement (change in position) per unit time h (vector)

$$\vec{v}(t + h/2) = \frac{\vec{r}(t + h) - \vec{r}(t)}{h}$$

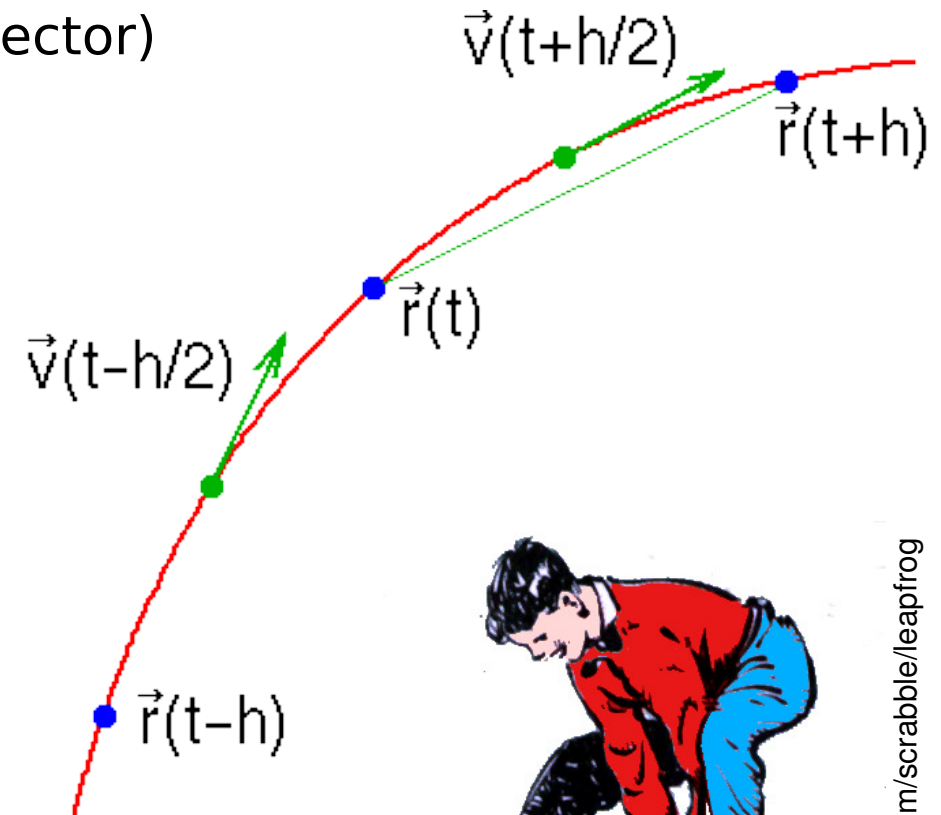
acceleration = change in velocity per unit time

$$\vec{a}(t) = \frac{\vec{v}(t + h/2) - \vec{v}(t - h/2)}{h} = \frac{\vec{f}}{m}$$

⇒

$$\left. \begin{aligned} \vec{v}(t + h/2) &:= \vec{v}(t - h/2) + \vec{a}(t)h \\ \vec{r}(t + h) &:= \vec{r}(t) + \vec{v}(t + h/2)h \\ t &:= t + h \end{aligned} \right\} \text{repeated}$$

- equivalent to Verlet (identical trajectory)
but: velocities at different time, a bit different (by $\mathcal{O}(h^2)$) kinetic energy



Leap-frog:

$$\left. \begin{aligned} v(t + h/2) &:= v(t - h/2) + a(t)h \\ r(t + h) &:= r(t) + v(t + h/2)h \\ t &:= t + h \end{aligned} \right\} \text{repeated}$$

2nd equation twice in 2 different times:

$$\begin{aligned} r(t + h) &= r(t) + v(t + h/2)h && \times + 1 \\ r(t) &= r(t - h) + v(t - h/2)h && \times - 1 \end{aligned}$$

Subtract both equations:

$$r(t + h) - r(t) = r(t) - r(t - h) + v(t + h/2)h - v(t - h/2)h$$

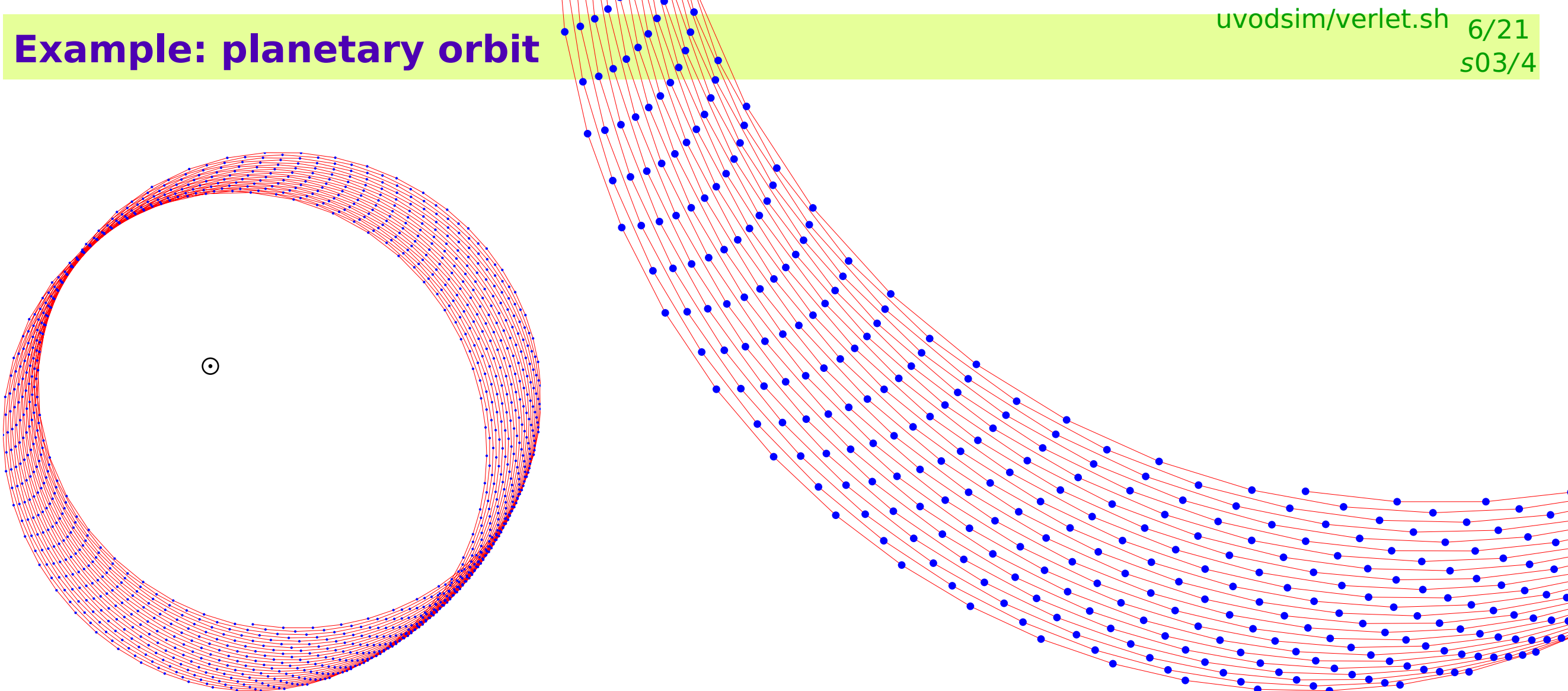
insert for the difference of velocities:

$$r(t + h) - 2r(t) + r(t - h) = h[v(t + h/2) - v(t - h/2)] = a(t)h^2 = \frac{f(t)}{m}h^2$$

which is the Verlet method

Example: planetary orbit

uvodsim/verlet.sh 6/21
s03/4



- energy is well conserved
- perihelion precession $\mathcal{O}(h^2)$
- harmonic oscillator: frequency shifted $\mathcal{O}(h^2)$

Euler–Lagrange equations

Our world: $\vec{r}^N = \{\vec{r}_1, \dots, \vec{r}_N\}$, $\dot{\vec{r}}^N = \{\dot{\vec{r}}_1, \dots, \dot{\vec{r}}_N\}$

Function $\mathcal{L} = \mathcal{L}(\vec{r}^N, \dot{\vec{r}}^N)$

Action:

$$S = \int_{t_0}^{t_1} \mathcal{L} dt$$

is stationary (likely min or max) between fixed points $\vec{r}^N(t_0)$ and $\vec{r}^N(t_1)$ for

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\vec{r}}_i} = \frac{\partial \mathcal{L}}{\partial \vec{r}_i}$$

Total $3N$ equations.

If \mathcal{L} = Lagrangian, then this is the **Hamilton principle**, or (in general) the “principle of minimum action” or so.

$$S = \int_{t_0}^{t_1} \mathcal{L} dt$$

Trajectory **variation**:

$$\vec{r}^N(t) \rightarrow \vec{r}^N(t) + \delta \vec{r}^N(t), \quad \delta \vec{r}^N(t_0) = \delta \vec{r}^N(t_1) = 0$$

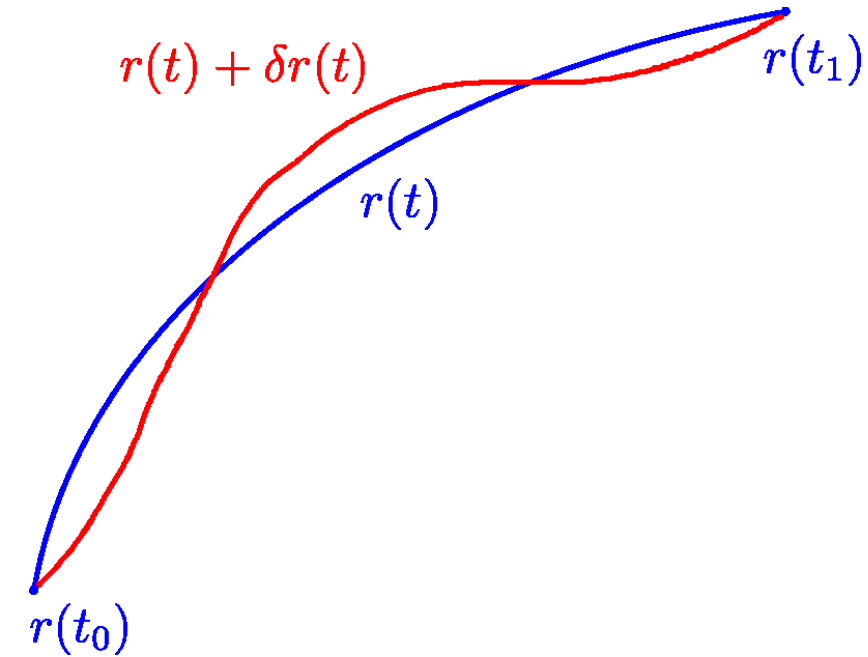
$$\delta S = \int_{t_0}^{t_1} \sum_i \frac{\partial \mathcal{L}}{\partial \vec{r}_i} \cdot \delta \vec{r}_i dt + \int_{t_0}^{t_1} \sum_i \frac{\partial \mathcal{L}}{\partial \dot{\vec{r}}_i} \cdot \delta \dot{\vec{r}}_i dt$$

The 2nd term integrated *by parts*:

$$\delta S = \left[\sum_i \frac{\partial \mathcal{L}}{\partial \dot{\vec{r}}_i} \cdot \delta \vec{r}_i \right]_{t_0}^{t_1} + \int_{t_0}^{t_1} \sum_i \delta \vec{r}_i \cdot \left[\frac{\partial \mathcal{L}}{\partial \vec{r}_i} - \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\vec{r}}_i} \right] dt$$

(1st [] = 0 because the endpoints are fixed)

$\delta \vec{r}_i$ are arbitrary \Rightarrow 2nd [] = 0



Math refreshment: Legendre transform

+ 9/21
s03/4

Let us have $f(x)$, better a convex one.

$$f^* = f - x \frac{df}{dx} \text{ "as function of } p = \frac{df}{dx}\text{"}$$

In a more mathematical language:

$$f^*(p) = \min_x (f - xp)$$

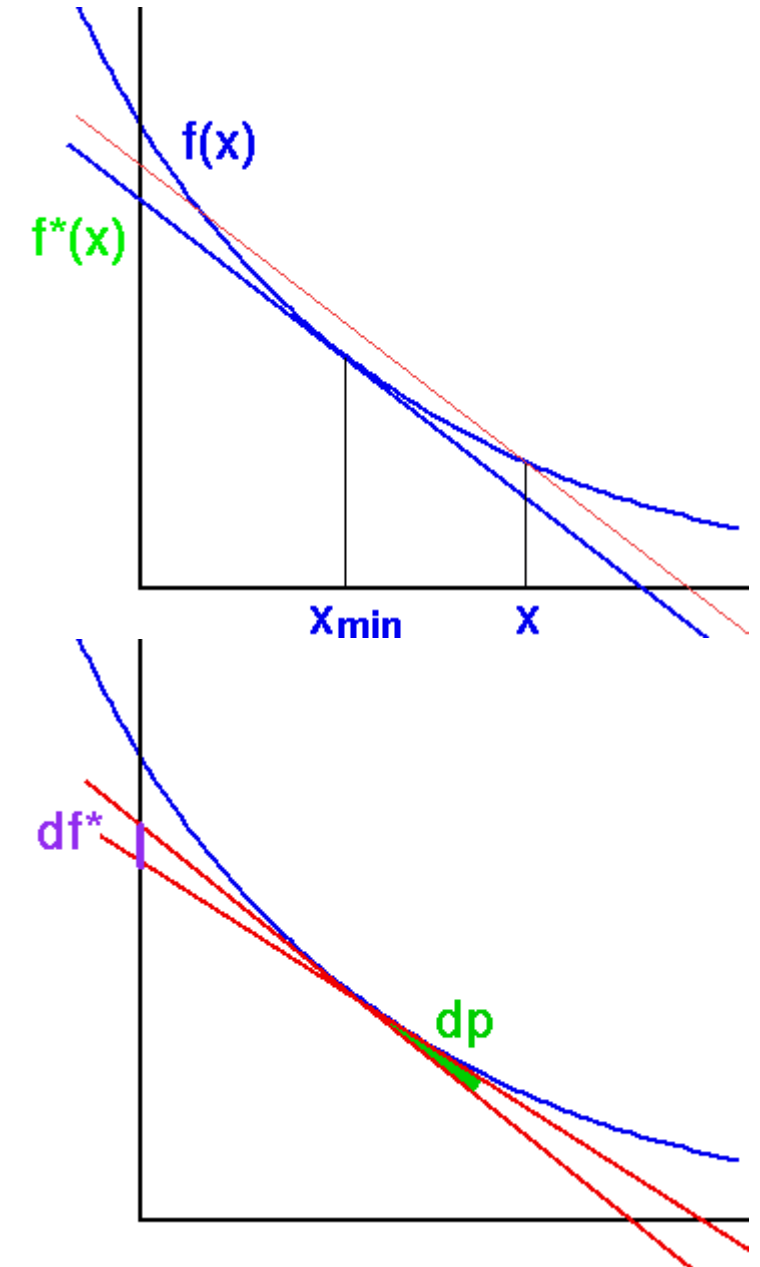
Differentials:

$$df = \frac{df}{dx} dx = p dx$$

$$df^* = df - d(px) = p dx - p dx - x dp = -x dp$$

And the reverse transformation:

$$\frac{df^*}{dp} = -x, \quad f^{**} = f^* - \frac{df^*}{dp} p = f^* + px = f$$



Internal energy $U = U(S, V)$:

$$dU = -p dV \quad [\text{ad.}]$$

$U(V)$ [ad.] is **convex**, because $p = -\frac{\partial U}{\partial V}$ is a **decreasing** function of V

Enthalpy $H = H(S, p)$:

$$H = U - \frac{\partial U}{\partial V} V = U + pV$$

$$dH = V dp \quad [\text{ad.}]$$

Reversed:

$$U = H - Vp = H - \frac{\partial H}{\partial p} p$$

Similarly $U(S) \rightarrow F(T)$, $F(N) \rightarrow \Omega(\mu)$, ...

Example. Plot $F(V)$ and $G(V)$ at constant T for the van der Waals equation of state
control: $\alpha \rightarrow T$

Let

$$\mathcal{L} = \mathcal{L}(\vec{r}_i^N, \dot{\vec{r}}_i^N) = E_{\text{kin}} - E_{\text{pot}} = \sum_i \frac{1}{2} m_i \dot{\vec{r}}_i^2 - U(\vec{r}_i^N)$$

then Lagrange equations = Newton's equations

Ummm ... nothing new yet.

But in the **generalized coordinates**

$$q_j = q_j(\vec{r}_1, \dots, \vec{r}^N), \quad j = 1 \dots 3N$$

it works, too!

Example: planet in the polar coordinates (r, ϕ)

$$\mathcal{L} = E_{\text{kin}} - E_{\text{pot}} = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\phi}^2) + \frac{K}{r}$$

Euler-Lagrange equations:

$$m\ddot{r} = mr\dot{\phi}^2 - \frac{K}{r^2} \quad (\text{Verlet not applicable})$$

$$mr^2\ddot{\phi} = 0 \quad \Rightarrow \quad mr^2\dot{\phi} = \text{const} \quad (\text{angular momentum})$$

From Lagrange to Hamilton

+ 12/21
s03/4

Momentum $\vec{p}_i = m_i \dot{\vec{r}}_i = \frac{\partial \mathcal{L}}{\partial \dot{\vec{r}}_i}$

Generalized momenta (definition): $p_j = \frac{\partial \mathcal{L}}{\partial \dot{q}_j}$

Example (planet): $p_\phi = mr^2 \dot{\phi}$

Legendre transform: $\dot{\vec{r}}_i \rightarrow \vec{p}_i$ (and opposite sign)

$$\mathcal{H} = \mathcal{H}(\vec{r}^N, \vec{p}^N) = \sum_i \vec{p}_i \cdot \dot{\vec{r}}_i - \mathcal{L}$$

$$\mathcal{L} = E_{\text{kin}} - E_{\text{pot}}$$

\mathcal{H} is called the **Hamiltonian**

Cartesian coordinates: $\mathcal{H} = E_{\text{kin}} + E_{\text{pot}}$

Using the Lagrange equations: $\dot{\vec{p}}_i = \frac{\partial \mathcal{L}}{\partial \vec{r}_i}$

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\vec{r}}_i} = \frac{\partial \mathcal{L}}{\partial \vec{r}_i}$$

⇒ **Hamilton's equations:**

$$\dot{\vec{p}}_i = -\frac{\partial \mathcal{H}}{\partial \vec{r}_i}, \quad \dot{\vec{r}}_i = \frac{\partial \mathcal{H}}{\partial \vec{p}_i}$$

Conservation of energy

+ 13/21
s03/4

Change of \mathcal{L} if both positions and velocities change
(not time: E_{pot} is assumed to be conservative $\Rightarrow \frac{\partial \mathcal{L}}{\partial t} = 0$)

$$d\mathcal{L} = \sum_i \left[\frac{\partial \mathcal{L}}{\partial \vec{r}_i} \cdot d\vec{r}_i + \frac{\partial \mathcal{L}}{\partial \dot{\vec{r}}_i} \cdot d\dot{\vec{r}}_i \right] = \sum_i (\dot{\vec{p}}_i \cdot d\vec{r}_i + \vec{p}_i \cdot d\dot{\vec{r}}_i)$$

Legendre transform:

$$d\mathcal{H} = \sum_i d(\vec{p}_i \cdot \dot{\vec{r}}_i) - d\mathcal{L} = \sum_i [-\dot{\vec{p}}_i \cdot d\vec{r}_i + \dot{\vec{r}}_i d\vec{p}_i] \stackrel{!}{=} \sum_i \left[\frac{\partial \mathcal{H}}{\partial \vec{r}_i} \cdot d\vec{r}_i + \frac{\partial \mathcal{H}}{\partial \vec{p}_i} \cdot d\vec{p}_i \right]$$

Hamilton equations:

$$\dot{\vec{p}}_i = -\frac{\partial \mathcal{H}}{\partial \vec{r}_i}, \quad \dot{\vec{r}}_i = \frac{\partial \mathcal{H}}{\partial \vec{p}_i}$$

And also:

$$\frac{d\mathcal{H}}{dt} = \sum_i \left[\frac{\partial \mathcal{H}}{\partial \vec{r}_i} \cdot \dot{\vec{r}}_i + \frac{\partial \mathcal{H}}{\partial \vec{p}_i} \cdot \dot{\vec{p}}_i \right] = 0$$

= conservation of energy (Hamiltonian = integral of motion)

A sledgehammer not needed to crack a nut

+ 14/21
s03/4

$$\begin{aligned}\frac{d}{dt}(E_{\text{kin}} + E_{\text{pot}}) &= \frac{d}{dt} \left[\sum_i \frac{m_i}{2} \dot{\vec{r}}_i^2 + U(\vec{r}^N) \right] \\ &= \sum_i \left[m_i \dot{\vec{r}}_i \cdot \ddot{\vec{r}}_i + \frac{\partial U}{\partial \vec{r}_i} \cdot \dot{\vec{r}}_i \right] = \sum_i \dot{\vec{r}}_i \cdot [m_i \ddot{\vec{r}}_i - \vec{f}_i] = 0\end{aligned}$$



Any (differentiable) symmetry (of the action) of a physical system has a corresponding conservation law.

- Time → energy conservation
(assuming $E_{\text{pot}}(t) = E_{\text{pot}}(t + \delta t)$)

- Translation → momentum

$$U(\vec{r}^N + \delta \vec{r}) = U(\vec{r}^N) \Rightarrow 0 = \delta \vec{r} \cdot \sum_i \frac{\partial U}{\partial \vec{r}_i} = -\delta \vec{r} \cdot \frac{d}{dt} \sum_i m_i \dot{\vec{r}}_i$$

Since $\delta \vec{r}$ is arbitrary, **total momentum** is conserved

- Rotation → angular momentum

$$\begin{aligned} U(\vec{r}^N + \delta \vec{\alpha} \times \vec{r}^N) &= U(\vec{r}^N) \\ \Rightarrow 0 &= \sum_i (\delta \vec{\alpha} \times \vec{r}_i) \cdot \frac{\partial U}{\partial \vec{r}_i} = - \sum_i (\delta \vec{\alpha} \times \vec{r}_i) \cdot m_i \ddot{\vec{r}}_i \\ &= - \sum_i \delta \vec{\alpha} \cdot (\vec{r}_i \times m_i \ddot{\vec{r}}_i) = -\delta \vec{\alpha} \cdot \frac{d}{dt} \sum_i \vec{r}_i \times m_i \dot{\vec{r}}_i \end{aligned}$$



(Amalie) Emmy
Noether

credit: Wikipedia

**Total angular
momentum**
is conserved

Let $f = f(\vec{r}^N, \vec{p}^N)$. Time development: $f(t + dt) = f(t) + \dot{f}dt$.

$$\frac{df}{dt} \equiv \dot{f} = \sum_i \left[\dot{\vec{r}}_i \cdot \frac{\partial f}{\partial \vec{r}_i} + \dot{\vec{p}}_i \cdot \frac{\partial f}{\partial \vec{p}_i} \right] = \sum_i \left[\frac{\partial \mathcal{H}}{\partial \vec{p}_i} \cdot \frac{\partial f}{\partial \vec{r}_i} - \frac{\partial \mathcal{H}}{\partial \vec{r}_i} \cdot \frac{\partial f}{\partial \vec{p}_i} \right] \equiv \{f, \mathcal{H}\}$$

$\{, \}$ is called the **Poisson bracket**

It holds $\{A, B\} = -\{B, A\}$

If $f = f(\vec{r}^N, \vec{p}^N)$ is an integral of motion, then $\{f, \mathcal{H}\} = 0$.

If $f = f(\vec{r}^N, \vec{p}^N, t)$ is an integral of motion, then $\{f, \mathcal{H}\} + \frac{\partial f}{\partial t} = 0$

Let us define the **Liouville operator**

$$i\hat{L} = \sum_i \left[\dot{\vec{r}}_i \cdot \frac{\partial}{\partial \vec{r}_i} + \dot{\vec{p}}_i \cdot \frac{\partial}{\partial \vec{p}_i} \right] = \sum_i \left[\frac{\partial \mathcal{H}}{\partial \vec{p}_i} \cdot \frac{\partial}{\partial \vec{r}_i} - \frac{\partial \mathcal{H}}{\partial \vec{r}_i} \cdot \frac{\partial}{\partial \vec{p}_i} \right] \equiv i\hat{L}_r + i\hat{L}_p$$

then (for $\frac{\partial f}{\partial t} = 0$)

$$\dot{f} = \{f, \mathcal{H}\} = i\hat{L}f$$

Postulate: $\{, \} \rightarrow i\hbar[,]$ signs wrong – see Czech version!

E.g.: $\{p, x\} = -1 \Rightarrow [\hat{p}, \hat{x}] = -i\hbar$

(x, p = any pair of conjugate canonical variables)

x-representation: $\psi = \psi(x)$, $\hat{x} = x$, $\hat{p} = -i\hbar \frac{\partial}{\partial x}$

In other words $[-i\hbar \frac{\partial}{\partial x}, x]\psi = -i\hbar\psi$ (well-known)

Test of the machinery: $\{p, f\} = -\frac{\partial f}{\partial x} \rightarrow [-i\hbar \frac{\partial}{\partial x}, f]\psi = -i\hbar \frac{\partial f}{\partial x} \psi$

Similarly for $f = f(\vec{r}^N, \vec{p}^N, t)$:

$$\{f, \mathcal{H}\} = \frac{\partial f}{\partial t} \rightarrow [f, \hat{H}] = i\hbar \frac{\partial f}{\partial t} \quad \text{i.e.,} \quad [f, \hat{H}]\psi = i\hbar \frac{\partial f}{\partial t} \psi$$

Satisfied by $\hat{H} = i\hbar \frac{\partial}{\partial t}$ (time Schrödinger equation); we write it as

$$\hat{H}\psi = i\hbar \frac{d}{dt} \psi$$

= time development of $\psi(x)$ (x cannot depend on time)

$$\dot{f} = i\hat{L}f$$

Formal (operator) solution (separation of variables)

$$\ln f = i\hat{L}t, \quad f(t) = \exp(i\hat{L}t) = \lim_{n \rightarrow \infty} (1 + i\hat{L}t/n)^n$$

What does this mean?

● consecutively $n \times$ repeated (approximately)

$$f(0 + t/n) = (1 + i\hat{L}t/n)f(0) = f(0) + \left. \frac{df}{dt} \right|_{t=0} t/n$$

● Taylor:

$$\begin{aligned} \exp(i\hat{L}t)f(0) &= 1 + (i\hat{L}f)t + (i\hat{L}i\hat{L}f)\frac{t^2}{2} + \dots = \\ &= 1 + \dot{f}(0)t + \ddot{f}\frac{t^2}{2} + \dots = f(t) \end{aligned}$$

The same Taylor-like trick for $i\hat{L}_r$ and $i\hat{L}_p$:

$$\begin{aligned}\exp(i\hat{L}_r t)f(\vec{r}^N, \vec{p}^N) &= 1 + (i\hat{L}_r f)t + (i\hat{L}_r i\hat{L}_r f)\frac{t^2}{2} + \dots = \\ &= 1 + \sum_i \dot{\vec{r}}_i \cdot \frac{\partial f}{\partial \vec{r}_i} t + \sum_j \dot{\vec{r}}_j \sum_i \dot{\vec{r}}_i : \frac{\partial^2 f}{\partial \vec{r}_i \partial \vec{r}_j} \frac{t^2}{2} + \dots = f(\vec{r}^N + \dot{\vec{r}}^N t, \vec{p}^N)\end{aligned}$$

$$\exp(i\hat{L}_p t)f(\vec{r}^N, \vec{p}^N) = f(\vec{r}^N, \vec{p}^N + \dot{\vec{p}}^N t)$$

Problem: operators $i\hat{L}_p$ and $i\hat{L}_r$ do not commute:

$$\exp(i\hat{L}) = \exp(i\hat{L}_p + i\hat{L}_r) \neq \exp(i\hat{L}_p) \exp(i\hat{L}_r)$$

So let at least approximately (for small h), but always **reversibly**:

$$\exp(i\hat{L}h) \approx \exp(i\hat{L}_p h/2) \exp(i\hat{L}_r h) \exp(i\hat{L}_p h/2)$$

Step by step (N omitted):

$$\begin{aligned} & (\vec{p}(0) , \vec{r}(0)) \\ & (\vec{p}(0) + \dot{\vec{p}}(0)h/2 , \vec{r}(0)) \\ & (\vec{p}(0) + \dot{\vec{p}}(0)h/2 , \vec{r}(0) + (1/m)[\vec{p}(0) + \dot{\vec{p}}(0)h/2]h) \\ & (\vec{p}(0) + [\dot{\vec{p}}(0) + \dot{\vec{p}}(h)]h/2 , \vec{r}(0) + (1/m)[\vec{p}(0) + \dot{\vec{p}}(0)h/2]h) \end{aligned}$$

This is the so called **velocity Verlet**:

$$\begin{aligned} r(t+h) &= r(t) + v(t)h + \frac{f(t)h^2}{m} \frac{1}{2} \\ v(t+h) &= v(t) + \frac{f(t) + f(t+h)h}{m} \frac{1}{2} \end{aligned}$$

The same trajectory as Verlet with $v(t) = \frac{r(t+h) - r(t-h)}{2h}$

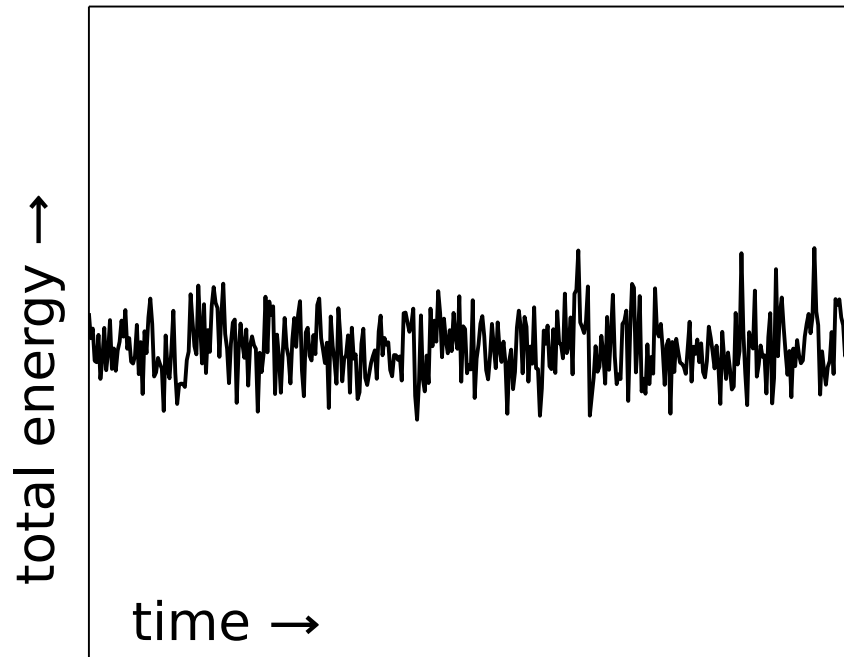
What is this good for?

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s03/4

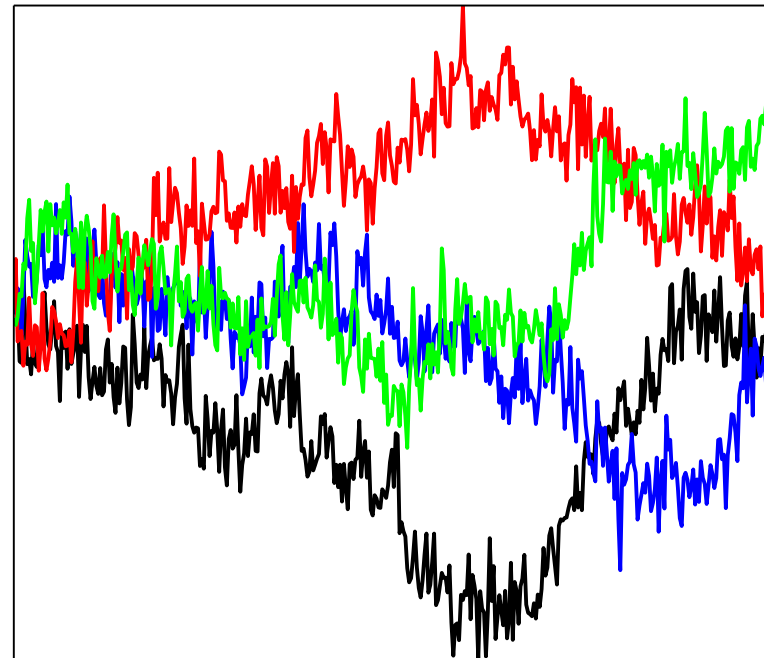
$$\exp(i\hat{L}_p h/2) \exp(i\hat{L}_r h) \exp(i\hat{L}_p h/2) = \exp(i\hat{L} h + \epsilon)$$

- error ϵ can be estimated ($\propto h^3$)
- we can calculate a “perturbed Hamiltonian” (error $\mathcal{O}(h^3)$ per step, $\mathcal{O}(h^2)$ globally), exactly constant with the Verlet method
i.e., Verlet is **symplectic** \Rightarrow error is bound
(time reversibility \Rightarrow only error $\propto t^{1/2}$)
- multiple-timestep methods and higher-order methods

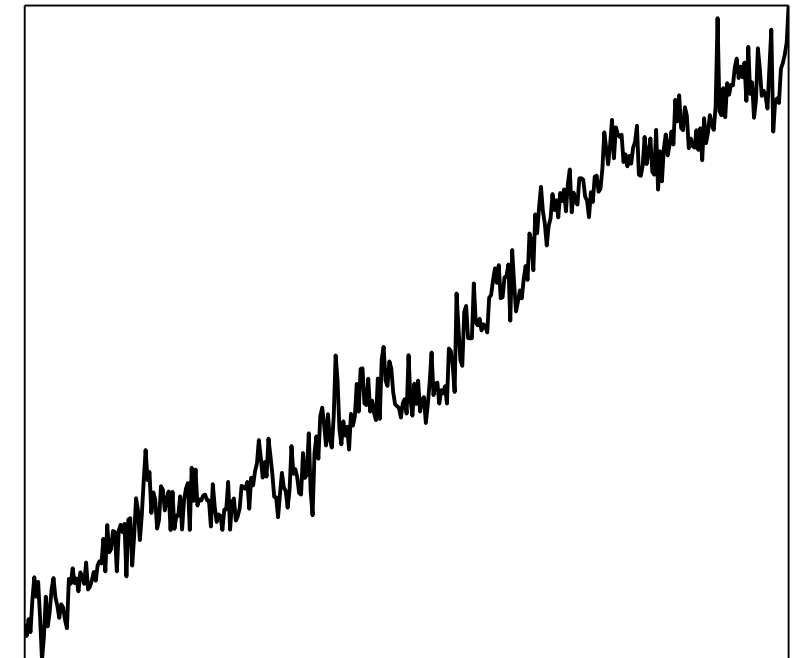
energy conservation error is used to set the timestep h



symplectic



reversible



irreversible